# Open topics in applied mathematics: Fast Methods in Scientific Computation

### MAT 393 C

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(These notes will be posted on the class webpage.)

### **Purpose of class:**

- The central theme is "fast" methods for solving elliptic PDEs such as:
  - The Laplace and Poisson equations.
  - Helmholtz' equation.
  - Time-harmonic Maxwell's equation.
  - The equations of linear elasticity.
  - The Stokes equation.
- We will also cover other computational methods, including:
  - FFT and other expansion based fast solvers.
  - Fast methods for *N*-body problems such as the Fast Multipole Method (FMM).
  - Techniques for accelerating matrix computations randomized methods for factorizing matrices, sparse solvers, Krylov methods, "rank-structured" matrix computations, etc.
- Light emphasis on proofs and theory. Stronger emphasis on practical computing.
- Focus is on numerical methods and scientific computing, but connections to applications will be discussed as well.

### **Definition of the term "fast":**

We say that a numerical method is *fast* if its execution time scales as O(N) as the problem size N grows.

Methods whose complexity is  $O(N \log N)$  or  $O(N \log^2 N)$  are also called "fast".





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Using a method that scales as O(N), problems that are 1000 times larger can be solved.

**Caveat:** It appears that Moore's law is no longer operative.

Processor speed is currently increasing quite slowly.

The principal increase in computing power is coming from *parallelization*.

In consequence, successful algorithms must scale well both with problem size and with the number of processors that a computer has.

To slightly offset the difficulty of parallelization, the *cost of storage is decreasing*. However, the speed of access is increasing only slowly, again reinforcing the need to keep data local in designing algorithms.

# Laplace's equation (in two dimensions for simplicity)

Let  $u = u(\mathbf{x})$  denote a differentiable function of the vector valued variable  $\mathbf{x} = (x_1, x_2)$ . The *Laplace operator* is defined by

$$\Delta u = \frac{\partial^2 u}{\partial x_1^2} + \frac{\partial^2 u}{\partial x_2^2}.$$

Let  $\Omega$  denote a *domain* with boundary  $\Gamma$ . Then the *Poisson equation* on  $\Omega$  is

(1) 
$$\begin{cases} -\Delta u(\boldsymbol{x}) = f(\boldsymbol{x}), & \boldsymbol{x} \in \Omega, \\ u(\boldsymbol{x}) = g(\boldsymbol{x}), & \boldsymbol{x} \in \Gamma. \end{cases}$$

The function *f* is a given *body load* and *g* is a given *boundary data*. If f = 0, we call (1) the *Laplace equation*.

The Poisson and Laplace equations are the simplest equations in a large class of so called *elliptic PDEs*. Other examples include Helmholtz, elasticity, Maxwell (for the "time-harmonic case").

### The Laplace and Poisson equations:

## **Electrostatics:**

$$egin{aligned} -\Delta\,u(oldsymbol{x}) &= f(oldsymbol{x}), &oldsymbol{x} \in \Omega, \ u(oldsymbol{x}) &= g(oldsymbol{x}), &oldsymbol{x} \in \Gamma. \end{aligned}$$

*u* is the electric potential

*f* is the electric charge density

g is a fixed potential on the boundary (Neumann b.c.  $\Rightarrow$  fixed fluxes)

Examples of applications:

- Design of electric engines / turbines / etc.
- Biochemical modeling.
- Design of electronic circuits.

("Magnetostatics" is entirely analogous.)

The *Laplace and Poisson* equations:

# **Gravity:**

$$-\Delta u(\mathbf{x}) = f(\mathbf{x}), \qquad \mathbf{x} \in \mathbb{R}^3,$$

*u* is the gravitational potential *f* is the mass density

Examples of applications:

• Astrophysics

*A "hidden" Laplace problem:* Consider a situation with *N* gravitational bodies in  $\mathbb{R}^3$ . Each body has mass  $m_i$  and location  $x_i$ . Then the force on body *i* resulting from interactions with the other bodies is

$$m{F}_i = \sum_{j 
eq i} G \, m_i \, m_j \, rac{m{x}_i - m{x}_j}{|m{x}_i - m{x}_j|^3},$$

where  $G \approx 6.67428 \cdot 10^{-11} \mathrm{m}^3/(\mathrm{kg\,s}^2)$  is the gravitational constant.

We now observe that the force  $F_i$  can be expressed as

$$\mathbf{F}_i = -m_i \sum_{j \neq i} \nabla u_j(\mathbf{x}_i),$$

where  $u_j = u_j(\mathbf{x})$  is the gravitational potential generated by the *j*'th charge

$$u_j(\boldsymbol{x}) = \sum_{j \neq i} G m_j rac{1}{|\boldsymbol{x} - \boldsymbol{x}_j|}$$

The potential  $u_i$  satisfies

$$-\Delta u_j(\mathbf{x}) = m_j \, \delta(\mathbf{x} - \mathbf{x}_j)$$

The total field  $u = \sum_{i} u_{i}$  satisfies

$$-\Delta u(\mathbf{x}) = f(\mathbf{x}) = \sum_{j} m_{j} \, \delta(\mathbf{x} - \mathbf{x}_{j}).$$

The problem of computing a sum such as

$$m{F}_i = \sum_{j 
eq i} G m_i m_j rac{m{x}_i - m{x}_j}{|m{x}_i - m{x}_j|^3},$$

arises directly in many applications:

- Astrophysics.
- Biochemical simulations (each "particle" is a charged part of a molecule).
- Modeling of semi-conductors (each "particle" is an ion).
- Fluid dynamics (each "particle" is an "vortex").

It also arises indirectly in many "fast" methods for solving elliptic PDEs.

The naïve computation of  $\{F_i\}_{i=1}^N$  requires  $O(N^2)$  operations since there are N(N-1)/2 "pair-wise interactions."

We will study in some detail a method that requires only O(N) operations; the so called *Fast Multipole Method* or *FMM*.

The *Laplace and Poisson* equations:

**Thermostatics:** 

$$egin{aligned} & -\Delta\,u(oldsymbol{x}) = f(oldsymbol{x}), & oldsymbol{x} \in \Omega, \ & u(oldsymbol{x}) = g(oldsymbol{x}), & oldsymbol{x} \in \Gamma. \end{aligned}$$

*u* is the temperature

*f* is the heat source density

g is a fixed temperature on the boundary (Neumann b.c.  $\Rightarrow$  fixed flows)

Examples of applications:

• ...

The *Helmholtz* equation:

Recall the *wave equation*:

$$\frac{\partial^2 u}{\partial x_1^2} + \frac{\partial^2 u}{\partial x_2^2} = \frac{\partial^2 u}{\partial t^2}.$$

The wave equation models vibrations in membranes, acoustic waves, certain electro-magnetic waves, and many other phenomena.

Now assume that the time dependence is "time harmonic":

$$u(\boldsymbol{x},t) = v(\boldsymbol{x}) \cos(\omega t).$$

Then  $\frac{\partial^2 u}{\partial t^2} = -\omega^2 u$  and so the wave equation becomes the *Helmholtz equation*:

$$\frac{\partial^2 v}{\partial x_1^2} + \frac{\partial^2 v}{\partial x_2^2} = -\omega^2 v.$$

#### The Maxwell equations

$$\begin{cases} \nabla \cdot \mathbf{E} = \rho & \nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} \\ \nabla \cdot \mathbf{B} = \mathbf{0} & \nabla \times \mathbf{B} = \mathbf{J} + \frac{\partial \mathbf{E}}{\partial t} \end{cases}$$

model electromagnetism. **E** is the electric field, and **B** is the magnetic field.

Consider the stationary case where  $\partial \mathbf{E}/\partial t = 0$  and  $\partial \mathbf{B}/\partial t = 0$ . Since **E** is curl-free, there exists a function  $u = u(\mathbf{x})$  such that

$$\mathbf{E} = -\nabla u.$$

(The function *u* is the electric potential.) We now find that

$$\rho = \nabla \cdot \mathbf{E} = \nabla \cdot (-\nabla u) = -\Delta u,$$

and we recover the Poisson equation we saw earlier:

$$-\Delta u = \rho.$$

#### The Maxwell equations

$$\begin{cases} \nabla \cdot \mathbf{E} = \rho & \nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} \\ \nabla \cdot \mathbf{B} = \mathbf{0} & \nabla \times \mathbf{B} = \mathbf{J} + \frac{\partial \mathbf{E}}{\partial t} \end{cases}$$

model electromagnetism. **E** is the electric field, and **B** is the magnetic field.

Now consider another simplification: the "time-harmonic" case where

$$\mathbf{E}(\mathbf{x},t) = \mathbf{E}(\mathbf{x}) e^{i\omega t}, \qquad \mathbf{B}(\mathbf{x},t) = \mathbf{B}(\mathbf{x}) e^{i\omega t}.$$

Then

$$\frac{\partial}{\partial t}\mathbf{E} = i\omega\mathbf{E}, \quad \text{and} \quad \frac{\partial}{\partial t}\mathbf{B} = i\omega\mathbf{B}$$

Inserting these relations into the Maxwell equations, we obtain a system of "Helmholtz-like" equations

$$\begin{cases} \nabla \cdot \mathbf{E} = \rho & \nabla \times \mathbf{E} = -i\omega \mathbf{B} \\ \nabla \cdot \mathbf{B} = \mathbf{0} & \nabla \times \mathbf{B} = \mathbf{J} + i\omega \mathbf{E} \end{cases}$$

In special cases, the system simplifies to the plain Helmholtz equation ...

Recall:

$$\begin{cases} \nabla \cdot \mathbf{E} = \rho & \nabla \times \mathbf{E} = -i\omega \mathbf{B} \\ \nabla \cdot \mathbf{B} = \mathbf{0} & \nabla \times \mathbf{B} = \mathbf{J} + i\omega \mathbf{E} \end{cases}$$

Suppose  $\rho = 0$  and  $\mathbf{J} = \mathbf{0}$ . Then

$$\nabla \times \nabla \times \mathbf{E} = \nabla \times (-i\omega \mathbf{B}) = -i\omega(\nabla \times \mathbf{B}) = -i\omega(i\omega \mathbf{E}) = \omega^2 \mathbf{E}.$$

Now recall that for any vector field **F** we have

$$\nabla \times \nabla \times \mathbf{F} = \nabla (\nabla \cdot \mathbf{F}) - \Delta \mathbf{F}.$$

Consequently:

$$abla (
abla \cdot \mathbf{E}) - \Delta \mathbf{E} = \omega^2 \mathbf{E}.$$

Finally recall that  $\nabla \cdot \mathbf{E} = 0$  to obtain the "Helmholtz-like" equation

$$-\Delta \mathbf{E} = \omega^2 \mathbf{E}.$$

The equations of *linear elasticity* in  $\mathbb{R}^d$ :

$$\sum_{j,k,l=1}^{d} \frac{1}{2} E_{ijkl} \left( \frac{\partial^2 u_k}{\partial x_l \partial x_j} + \frac{\partial^2 u_l}{\partial x_k \partial x_j} \right) = f_i, \qquad i = 1, 2, \ldots, d.$$

The function  $\boldsymbol{u} = \boldsymbol{u}(\boldsymbol{x}) = (u_1(\boldsymbol{x}), u_2(\boldsymbol{x}), \dots, u_d(\boldsymbol{x}))$  is the displacement of an elastic material subjected to the body load  $\boldsymbol{f} = \boldsymbol{f}(\boldsymbol{x})$  at the point  $\boldsymbol{x}$ .

 $(E_{ijkl})_{i,j,k,l=1}^d$  is the stiffness tensor which describes the material properties.

Many simplifications can be derived from the basic equilibrium equation. For instance, if the material is isotropic, and if f = 0, then the displacements satisfy the *biharmonic equation* 

$$(-\Delta)^2 \boldsymbol{u} = 0.$$

Another simplification is the displacement of a thin elastic membrane:

$$egin{aligned} & ig(-\Delta)^2\,u(oldsymbol{x})=f(oldsymbol{x}), & oldsymbol{x}\in\Omega, \ & u(oldsymbol{x})=g(oldsymbol{x}), & oldsymbol{x}\in\Gamma, \ & u_n(oldsymbol{x})=h(oldsymbol{x}), & oldsymbol{x}\in\Gamma. \end{aligned}$$

Here *f* is the body load (e.g. gravity), *h* is the prescribed deflection at the boundary, and *h* is the prescribed normal derivative. (Since the equation has order *four*, we need *two* boundary conditions.)

### **Outline:**

Week:	Material covered:
1:	Introduction: Objectives of the course. Quick review of basic elliptic PDEs and their connec-
	tions to physical applications. Analytic solution formulas, and their relationship to numerical
	methods. Fast algorithms for global operators.
2:	Linear algebra: Review of basic matrix factorizations. Techniques for computing low-rank
	approximations to matrices. Randomized methods for matrix computations.
4:	Rank-structured matrices: What they are, where they arise in applications, how they enable
	fast solvers (and fast matrix algebra more generally).
5:	Krylov methods for solving linear systems and computing partial spectral decompositions.
7:	Fast solvers for elliptic PDEs based on the FFT and related techniques.
8:	Direct solvers for elliptic PDEs based on Gaussian elimination combined with nested dissec-
	tion ordering of the nodes ("multifrontal methods"). Sweeping solvers.
10:	Boundary integral equations. How a PDE can be rewritten as an integral equation. Advan-
	tages and disadvantages. Second kind Fredholm equations. Reduction of dimensionality.
12:	The Fast Multipole Method, and fast summation techniques. The kernel evaluation map.
	Kernel-independent FMMs and $\mathcal{H}$ -matrices.
14:	Fast direct solvers for integral equations.
15:	(If time permits) Johnson-and-Lindenstrauss theory, and connections to analysis of com-
	plex high dimensional data sets.

### **Practicalities:**

*Text:* There is no "official" text. The syllabus is defined by the material covered in class. Extensive latexed notes will be made available on the course website:

http://users.ices.utexas.edu/~pgm/Teaching/2019\_393C

Comments, errata, suggestions, ..., are highly appreciated!

Attendance: Strongly encouraged.

*Computer programming:* Matlab will be used. If you do not have access to a computer with Matlab, please contact the instructor.

*Grading:* No exam. Final grade is based on homeworks and a project:

- 50%: Five homework problems worth 10% each.
- 10%: Handing in a carefully latexed "reference solution".
- 40%: Final project.