

Linearly Stable Relative Equilibria Utilizing a Dominant Mass

N-BODY PROBLEM

Given n bodies in space with initial positions and velocities, under mutual gravitational attraction, what is their future motion? Define:

 $\mathbf{q}_i = \text{position of the i-th body in } \mathbb{R}^2$ $m_i = \text{mass of the i-th body}$

F = ma gives,

$$m_i \ddot{\mathbf{q}}_i = \sum_{j \neq i}^n \frac{m_i m_j (\mathbf{q}_j - \mathbf{q}_i)}{||\mathbf{q}_i - \mathbf{q}_j||^3}$$

for each $i \in (1, 2, ..., n)$

CENTRAL CONFIGURATIONS

A central configuration (c.c.) is an initial configuration of bodies \mathbf{x}_i , which satisfies the following equation

$$\sum_{j\neq i}^{n} \frac{m_i m_j (\mathbf{x}_j - \mathbf{x}_i)}{||\mathbf{x}_j - \mathbf{x}_i||^3} + \omega^2 m_i \mathbf{x}_i = 0 \qquad \forall i \in (1 \dots n)$$

for some proportionality constant ω^2 .

- Providing a c.c. with the correct choice of initial velocities leads to a periodic solution called a relative equilibrium.
- Released from rest, a c.c. maintains the same shape as it heads toward total collision.
- Any Kepler orbit can be attached to a c.c. to obtain a new solution to the *n*-body problem.
- **Key Fact**: Central configurations are critical points of the Newtonian potential function U restricted to the level surface $I = I_0$.

where
$$I = \frac{1}{2} \sum_{i=1}^{n} m_i ||\mathbf{x}_i||^2$$

LINEAR STABILITY

Due to symmetry, there are two invariant subspaces W_1 and W_2 that lead to the eight eigenvalues $0, 0, \pm \omega i, \pm \omega i, \pm \omega i$, for any relative equilibrium x. We say x is *nondegenerate* if the remaining 4n-8 eigenvalues are nonzero. It is *spectrally stable* if the eigenvalues are pure imaginary and is *linearly stable* if in addition, the restriction of the matrix obtained upon linearizing about x, to the skew-orthogonal complement of $W_1 \cup W_2$ is diagonalizable.

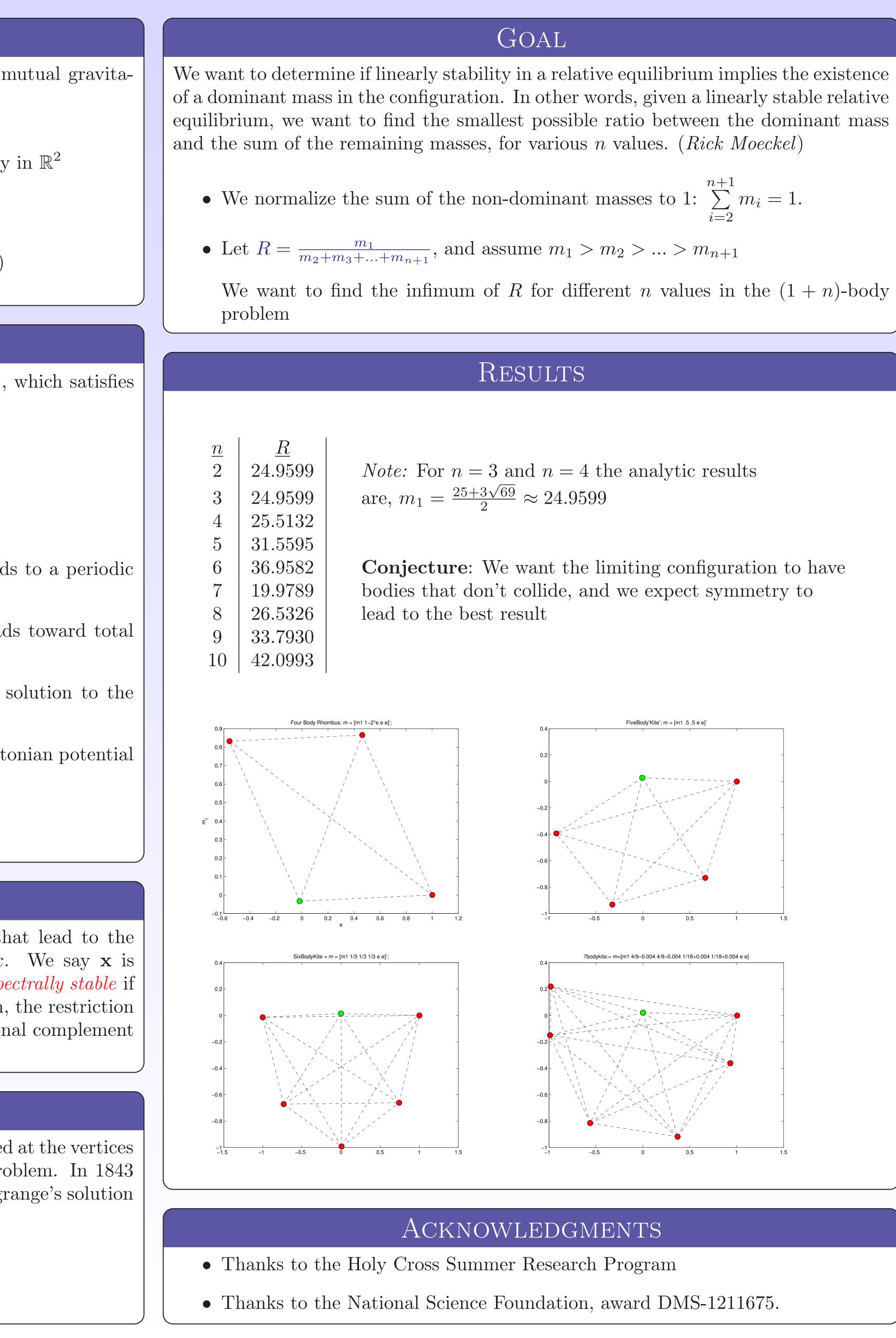
KNOWN RESULTS

• In 1772, Lagrange discovered that three bodies of any mass located at the vertices of an equilateral triangle represent a solution to the 3-body problem. In 1843 Gascheau (and later in 1875 Routh) determined stability for Lagrange's solution

$$\frac{m_1m_2 + m_1m_3 + m_2m_3}{(m_1 + m_2 + m_3)^2} < \frac{1}{27}$$

• All collinear relative equilibria are unstable

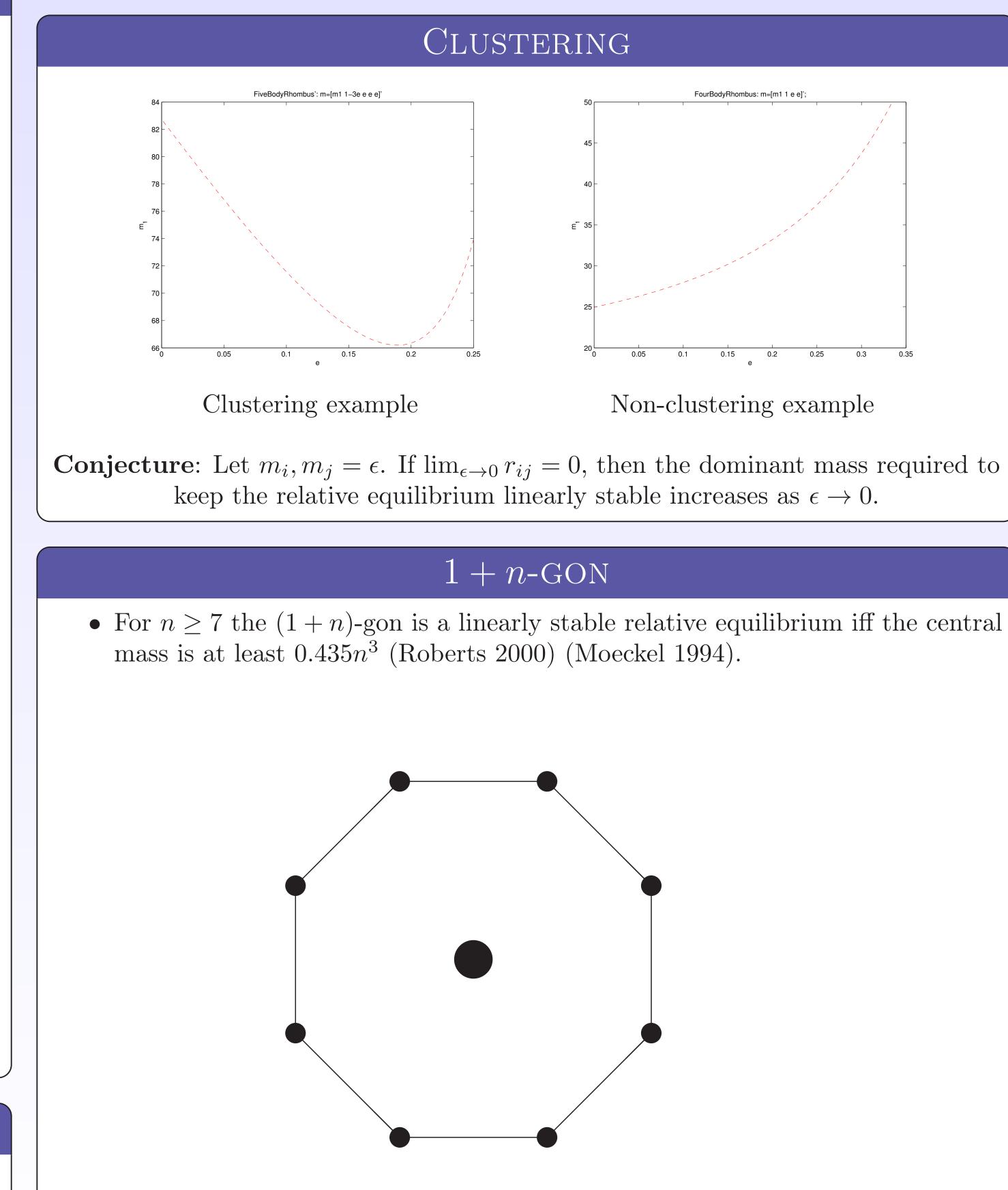
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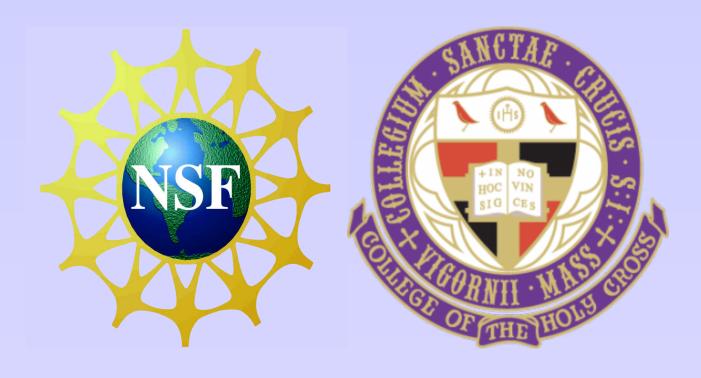


Lasses to 1:
$$\sum_{i=2}^{n+1} m_i = 1.$$

TECHNIQUES: MATLAB

- stability.
- Numerics: ues are 0 (within 1e-10).
- corresponding relative equilibrium is linearly stable.





• We use gradient flow to locate minima of the Newtonian potential function Uconstrained to a level surface of the moment of inertia $I = \frac{1}{2}$.

• Given an initial configuration \mathbf{x}_0 and a set of masses m, a MatLab function called config analyzer (x_0, m) was written to find a local minimum and determine its

ODE15s (stiff ODE), RelTol, AbsTol = 1e-10, and $||M^{-1}\nabla U(\mathbf{x}) + \omega^2 \mathbf{x}||$ is always $\leq 1e-13$. Check if real parts of eigenval-

• For each family of relative equilibria, we use a "binary search" method incorporating $config_analyzer()$ to locate the smallest possible m_1 such that the